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# Geometry of space-time and generalised Lagrange gauge theory

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Dedicated to Professor Lajos Tamássy on his 70th birthday

Abstract. In  $\S1$  and  $\S2$  the authors present the Einstein and Maxwell equations for the generalised Lagrange space

$$GL^n = (M, g_{ij}(x, y)) = e^{2\sigma(x, y)}\gamma_{ij}(x)),$$

and characterize the case of vanishing mixed curvature tensor field of the canonical linear d-connection. The Lagrangian gauge theory — in G.S. ASANOV's sense [1] is developed in §3 for the tangent bundle endowed with (h, v)-metrics, obtaining the generalised Einstein - Yang Mills equations with respect to the metric gauge tensor fields and to the gauge field  $\sigma(x, y)$  for three remarkable cases in which the metrics are derived from the fundamental tensor field  $g_{ij}(x, y)$ . Proofs are, in most cases, mechanical but rather tedious calculations. They are omitted.

#### Introduction

In a previous paper [6] it was shown that the EPS conditions can be satisfied in a more general frame than that of Finsler spaces, namely for convenient generalised Lagrange spaces. More precise, was examined the space  $GL^n = (M, g_{ij}(x, y))$ , where M is a *n*-dimensional differentiable manifold and

(A1) 
$$g_{ij}(x,y) = e^{2\sigma(x,y)}\gamma_{ij}(x),$$

 $\gamma_{ij}(x)$  being a metric tensor field on M, and  $\sigma: TM \to \mathbb{R}$  a function of class  $C^{\infty}$  on  $\widetilde{TM} = TM \setminus \{0\}$ , continuous on the null section of the tangent

bundle. The space was endowed with the non-linear connection

(A2) 
$$N^{i}{}_{j}(x,y) = \begin{cases} i\\ kj \end{cases} y^{k},$$

where  ${i \\ kj}$  are the Christoffel symbols for  $\gamma_{ij}(x)$ . Under these assumptions,  $GL^n$  represents a convenient relativistic model, since it has the same conformal and projective properties as the Riemannian space  $V^n = (M, \gamma_{ij}(x))$ .

## §1. Einstein equations for $GL^n$

Developing the formalism presented in [3,4], we remind that the canonical h- and v-symmetrical d-connection of  $GL^n$  has the coefficients

. .

(1.1) 
$$L_{jk}{}^{i} = \begin{cases} i\\ jk \end{cases} + \Lambda^{i}_{jk}$$
$$C_{bc}{}^{a} = \delta^{a}_{b}\dot{\sigma}_{c} + \delta^{a}_{c}\dot{\sigma}_{b} - \gamma_{bc}\dot{\sigma}^{a}$$

where we used the notations

(1.1') 
$$\begin{aligned} \Lambda^{i}_{jk} &= \delta^{i}_{j}\sigma_{k} + \delta^{i}_{k}\sigma_{j} - \gamma_{jk}\sigma^{i} \\ \sigma_{k} &= \delta_{k}\sigma, \ \dot{\sigma}_{a} = \dot{\partial}_{a}\sigma, \ \sigma^{k} = \gamma^{ks}\sigma_{s}, \ \dot{\sigma}^{a} = \gamma^{ab}\dot{\sigma}_{b} \end{aligned}$$

and

(1.1") 
$$\delta_k = \partial_k - N_k^a(x, y)\dot{\partial}_a, \quad \dot{\partial}_b = \frac{\partial}{\partial y^b}; \quad \partial_k \equiv \frac{\partial}{\partial x^k}.$$

Then the torsion d-tensor fields of the canonical linear d-connection  $C\Gamma(N)$ of  $GL^n$  have the coefficients given by

(1.2) 
$$\begin{cases} T_{jk}{}^{i} = 0, \quad S_{bc}^{a} = 0, \quad C_{ja}{}^{i}, \\ R_{kl}^{a} = r_{d}{}^{a}{}_{kl}y^{d}, \quad P_{kb}^{a} = -\Lambda_{bk}^{a}. \end{cases}$$

and the curvature d-tensor fields have the expressions

$$(1.3) \begin{cases} R_{jkl}^{\ i} = r_{jkl}^{\ i} + \delta^{i}_{(k}\sigma_{jl)} - \gamma^{is}\gamma_{j(k}\sigma_{sl)} + \gamma_{js}\dot{\sigma}^{(s}R_{kl}^{i)} \\ P_{j}^{\ i}_{\ kc} = \delta^{i}_{k}\overset{1}{\sigma}_{jc} - \delta^{i}_{c}\overset{2}{\sigma}_{jk} - \gamma^{is} \cdot \left[\gamma_{jk}\overset{1}{\sigma}_{sc} - \gamma_{jc}\overset{2}{\sigma}_{sk}\right] + \gamma^{is}\gamma_{ck}\sigma_{(s}\dot{\sigma}_{j)} \\ S^{a}_{bcd} = \delta^{a}_{(c}\dot{\sigma}_{bd)} - \gamma^{as}\gamma_{b(c}\dot{\sigma}_{sd)} \end{cases}$$

where  $r_j{}^i{}_{kl}$  is the curvature tensor field of  $\gamma_{ij}(x)$ ,  $t_{(ij)} = t_{ij} - t_{ji}$ , and we considered the following d-tensor fields

$$(1.3') \qquad \begin{aligned} \sigma_{sl} &= \sigma_{s|l} + \sigma_s \sigma_l - \frac{1}{2} \gamma_{sl}^{\ H} = \sigma_{s\uparrow l} - \sigma_s \sigma_l + \frac{1}{2} \gamma_{sl}^{\ H} \sigma_{sl}^{\ H} \\ \frac{1}{2} \sigma_{sl} &= \sigma_s|_l + \sigma_s \dot{\sigma}_l - \frac{1}{2} \gamma_{sl}^{\ M} \sigma_{sl}^{\ H} = \dot{\partial}_l \sigma_s - \dot{\sigma}_s \sigma_l + \frac{1}{2} \gamma_{sl}^{\ M} \sigma_{sl}^{\ H} \\ \frac{2}{2} \sigma_{sl} &= \dot{\sigma}_{s|l} + \dot{\sigma}_s \sigma_l - \frac{1}{2} \gamma_{sl}^{\ M} \sigma_{s\uparrow l}^{\ H} = \dot{\sigma}_{s\uparrow l} - \sigma_s \dot{\sigma}_l + \frac{1}{2} \gamma_{sl}^{\ M} \sigma_{sl}^{\ H} \\ \dot{\sigma}_{sl} &= \dot{\sigma}_s|_l + \dot{\sigma}_s \dot{\sigma}_l - \frac{1}{2} \gamma_{sl}^{\ V} \sigma_{s\uparrow l}^{\ H} = \dot{\partial}_l \dot{\sigma}_s - \dot{\sigma}_s \dot{\sigma}_l + \frac{1}{2} \gamma_{sl}^{\ V} \sigma_{s\downarrow l}^{\ H} \\ \dot{\sigma}_{sl} &= \dot{\sigma}_s|_l + \dot{\sigma}_s \dot{\sigma}_l - \frac{1}{2} \gamma_{sl}^{\ V} \sigma_{s\downarrow l}^{\ H} = \dot{\partial}_l \dot{\sigma}_s - \dot{\sigma}_s \dot{\sigma}_l + \frac{1}{2} \gamma_{sl}^{\ V} \sigma_{s\downarrow l}^{\ H} \\ \dot{\sigma}_{s\downarrow l} &= \dot{\sigma}_s|_l + \dot{\sigma}_s \dot{\sigma}_l - \frac{1}{2} \gamma_{s\downarrow l}^{\ V} \sigma_{s\downarrow l}^{\ H} \\ \dot{\sigma}_{s\downarrow l} &= \dot{\sigma}_s \dot{\sigma}_l + \frac{1}{2} \gamma_{s\downarrow l}^{\ V} \sigma_{s\downarrow l}^{\ H} \\ \dot{\sigma}_{s\downarrow l} &= \dot{\sigma}_s \dot{\sigma}_l + \frac{1}{2} \dot{\sigma}_{s\downarrow l}^{\ V} \sigma_{s\downarrow l}^{\ H} \\ \dot{\sigma}_{s\downarrow l} &= \dot{\sigma}_s \dot{\sigma}_l + \frac{1}{2} \dot{\sigma}_{s\downarrow l}^{\ V} \sigma_{s\downarrow l}^{\ H} \\ \dot{\sigma}_{s\downarrow l} &= \dot{\sigma}_s \dot{\sigma}_l + \frac{1}{2} \dot{\sigma}_{s\downarrow l}^{\ V} \sigma_{s\downarrow l}^{\ H} \\ \dot{\sigma}_{s\downarrow l} &= \dot{\sigma}_s \dot{\sigma}_l + \frac{1}{2} \dot{\sigma}_{s\downarrow l}^{\ V} \sigma_{s\downarrow l}^{\ V} \\ \dot{\sigma}_{s\downarrow l} &= \dot{\sigma}_s \dot{\sigma}_l + \frac{1}{2} \dot{\sigma}_{s\downarrow l}^{\ V} \sigma_{s\downarrow l}^{\ V} \\ \dot{\sigma}_{s\downarrow l} &= \dot{\sigma}_s \dot{\sigma}_l + \frac{1}{2} \dot{\sigma}_{s\downarrow l}^{\ V} \sigma_{s\downarrow l}^{\ V} \\ \dot{\sigma}_{s\downarrow l} &= \dot{\sigma}_s \dot{\sigma}_{s\downarrow l}^{\ V} \\ \dot{\sigma}_{s\downarrow l} &= \dot{\sigma}_s \dot{\sigma}_{s\downarrow$$

with  $\stackrel{H}{\sigma} = \sigma_s \sigma^s$ ,  $\stackrel{M}{\sigma} = \sigma_s \dot{\sigma}^s$ ,  $\stackrel{V}{\sigma} = \dot{\sigma}_s \dot{\sigma}^s$ . We also used the covariant derivatives

(1.3") 
$$\sigma_{s|l} = \delta_l \sigma_s - L_{sl}^t \sigma_t, \ \sigma_s|_l = \dot{\partial}_l \sigma_s - C_{sl}^t \sigma_t, \ \sigma_{s\uparrow l} = \partial_l \sigma_s - \begin{cases} t\\ sl \end{cases} \sigma_t.$$

Then, by contracting the indices, one can derive the four Ricci tensor fields, expressed by

(1.4)  

$$R_{ij} = r_{ij} - \gamma_{ij}\dot{\sigma} + (2-n)\cdot\sigma_{ij} + \gamma_{is}\dot{\sigma}^{(s}r_t^{a)}{}_{ja}$$

$$\frac{1}{P}_{bk} = (1-n)\hat{\sigma}_{bk} + \hat{\sigma}_{bk} - \gamma_{bk}\hat{\sigma} + \sigma_{(k}\dot{\sigma}_{b)}$$

$$\frac{2}{P}_{bc} = (1-n)\hat{\sigma}_{bc} - \hat{\sigma}_{bc} - \gamma_{bc}\hat{\sigma} + \sigma_{(c}\dot{\sigma}_{b)}$$

$$S_{bc} = (2-n)\dot{\sigma}_{bc} - \gamma_{bc}\dot{\sigma}$$

and then the curvature scalar fields

(1.5) 
$$\begin{cases} R = e^{-2\sigma} [r + 2(1-n)\overset{0}{\sigma} + 2r_{sc}y^s \dot{\sigma}^c] \\ \overset{1}{P} = -\overset{2}{P} = e^{-2\sigma} \cdot 2(1-n)\overset{1}{\sigma}, \quad S = e^{-2\sigma} \cdot 2(1-n)\dot{\sigma}, \end{cases}$$

where

$$\overset{0}{\sigma} = \gamma^{ij}\sigma_{ij}, \quad \overset{1}{\sigma} = \gamma^{ij}\overset{1}{\sigma}_{ij}, \quad \overset{2}{\sigma} = \gamma^{ij}\overset{2}{\sigma}_{ij}, \quad \dot{\sigma} = \gamma^{ab}\dot{\sigma}_{ab}, \quad r_{bk} = r^s_{bks}, \quad r = \gamma^{bk}r_{bk}.$$

Regarding the mixed curvature tensor field, the following result obtained by direct calculation holds true:

**Theorem.** The following assertions are valid:

a)  $P_{bkc}^{a} - P_{bck}^{a} = 0$  iff  $\overset{1}{\sigma}_{bc} + \overset{1}{\sigma}_{cb} = 0$ b)  $P_{bkc}^{a} + P_{bck}^{a} = 0$  iff  $\overset{1}{\sigma}_{bc} - \overset{1}{\sigma}_{cb} = \frac{2}{n}\sigma_{(b}\dot{\sigma}_{c)} = 0.$ c)  $P_{b}^{a}{}_{kc} = 0$  implies  $\overset{HV}{\sigma}\sigma = \overset{M2}{\sigma}$  and  $\overset{1}{\sigma} = \overset{2}{\sigma} = 0$ 

d) 
$$\stackrel{1}{P}_{bk} = \stackrel{2}{P}_{bk} = 0$$
 iff  $\stackrel{1}{\sigma}_{bk} = \frac{1}{n}\sigma_{(b}\dot{\sigma}_{k)}$   
e)  $\stackrel{1}{\sigma}_{jk} = \stackrel{2}{\sigma}_{kj}$ .

*Remark.* Using a), b), and e) it becomes obvious that  $P_b{}^a{}_{kc} = 0$  implies  ${}^1\sigma_{bk} = {}^2\sigma_{bk} = 0$ .

The Einstein equations of the space  $GL^n$  have the generic form

(1.6) 
$$\begin{cases} R_{ij} - \frac{1}{2}Rg_{ij} = \kappa T_{ij} \\ S_{ab} - \frac{1}{2}Sg_{ab} = \kappa T_{ab} \end{cases}$$

where  $\overset{H}{T}_{ij}$  and  $\overset{V}{T}_{ab}$  are called the *h*- and the *v*-components of the energymomentum tensor field of the space  $GL^n$  respectively,  $g_{ab} = \delta^i_a \delta^j_b \cdot g_{ij}$ and  $\kappa$  is the gravific constant. Using the expressions of the Ricci *d*-tensor fields (1.4) and of the curvature scalars (1.5) we obtain that the Einstein equations of the space  $GL^n$  admit the equivalent form

(1.6') 
$$\begin{cases} r_{ij} - \frac{1}{2}r\gamma_{ij} + t_{ij} = \kappa T_{ij} \\ (2-n)(\dot{\sigma}_{ab} - \dot{\sigma}\gamma_{ab}) = \kappa T_{ab} \end{cases}$$

where

$$t_{ij} = (n-2)(\gamma_{ij}\overset{0}{\sigma} - \sigma_{ij}) + \gamma_{ij}r_{st}y^s\dot{\sigma}^t + \gamma_{is}\dot{\sigma}^{(s}r^{a)}_{tja}y^t$$

We remark that in the first equation of the system (1.6') the term  $t_{ij}$  is complementary to the classical Einstein equations of the Riemannian space  $V^n = (M, \gamma_{ij}(x)).$ 

### $\S$ **2.** Maxwell equations for $GL^n$

Let the h- and v-deflection tensor fields be given respectively by

$$D^{i}{}_{j} = y^{i}{}_{|j}, \ d^{i}{}_{j} = y^{i}{}_{|j}$$

and the corresponding ones having the indices lowered

$$D_{ij} = g_{is} D^{s}{}_{j}, \ d_{ij} = g_{is} d^{s}{}_{j}.$$

Then we define the h- and v-electromagnetic tensor fields by

(2.1) 
$$F_{ij} = \frac{1}{2}(D_{ij} - D_{ji}), \quad f_{ij} = \frac{1}{2}(d_{ij} - d_{ji})$$

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respectively, and infer the Maxwell equations

(2.2) 
$$\begin{cases} \sigma F_{ij|k} = \sigma e^{2\sigma} r_0^{s} {}_{kj} y_i \dot{\sigma}_k \\ \sigma F_{ij|k} + \sigma f_{ij|k} = 0 \\ \sigma f_{ijk} f_{ij|k} = 0 \\ \sigma f_{ijk} f_{ij|k} = 0 \end{cases}$$

where  $\sigma_{ijk}$  denotes cyclic summation in the indices i, j, k and

$$r_0{}^s{}_{kj} = r_t{}^s{}_{kj}y^t, \ y_i = \gamma_{is}y^s.$$

Indeed, (2.2) takes place since  $D_{ij}$  and  $d_{ij}$  satisfy the relations

$$\begin{cases} D_{ij|k} - D_{ik|j} = R_{0ijk} - d_{is}R_{jk}^{s} \\ D_{ij}|_{k} - d_{ik|j} = P_{0ijk} - D_{is}C_{jk}^{s} - d_{is}P_{jk}^{s} \\ d_{ij}|_{k} - d_{ik}|_{j} = S_{0ijk} \end{cases}$$

and the Bianchi identities lead to (2.2). Also, the same equations come up clearly if we consider the expressions

$$F_{ij} = e^{2\sigma}(y_i\sigma_j - y_j\sigma_i), \quad f_{ij} = e^{2\sigma}(y_i\cdot\sigma_j - \gamma_j\sigma_i)$$

using direct computation.

In the following we examine the Maxwell equations for some remarkable particular examples of spaces  $GL^n$ :

1. If  $\sigma(x,y) = \frac{1}{2}\gamma_{rs}y^r y^s$ , then the Maxwell equations are trivially satisfied, since as  $\sigma_k = 0$  and  $\dot{\sigma}_k = y_k$ , we infer that

$$F_{ij} = f_{ij} = 0.$$

2. If  $GL^n$  is locally Minkowskian, i.e. at any point  $(x, y) \in TM$  there exists a domain of a local map in which

$$g_{ij}(x,y) = e^{2\sigma(y)}\gamma_{ij}$$
 with  $\partial_k\gamma_{ij} = 0$ ,

then  $F_{ij} = 0$  and  $f_{ij} = e^{2\sigma(y)}y_{(i}\dot{\sigma}_{j)}$ , and the Maxwell equations

$$\underset{ijk}{\sigma} f_{ij}|_k = 0$$

take place.

3. For a dispersive medium [7]  $(M, V^i(x), n(x, V(x)))$  where  $V^i(x)$  is the speed of the particle and n(x, V(x)) is the refraction index, we can consider

$$\sigma(x,y) = \alpha \cdot \left(1 - \frac{1}{n^2(x,y)}\right), \quad \alpha \in \mathbb{R}, \alpha > 0.$$

Remark that for  $y^i = V^i(x)$  the Maxwell equations have to be computed directly, being not implied by (2.2).

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### $\S3$ . Generalized Einstein - Yang Mills equations

Let TM be endowed with a (h, v)-metric  $G_*$  [3] given by

(3.1) 
$$G_* = g_{ij} dx^i \otimes dx^j + h_{ab} \delta y^a \otimes \delta y^b$$

where  $\{dx^i, \delta y^a \mid i, a = 1, n1, n\}$  is the adapted basis for the cotangent bundle, dual to (1.1"), and  $g_{ij}(x, y)$ ,  $h_{ab}(x, y)$  are symmetric *d*-tensor fields of rank n.

Let the coordinate transform on TM be given by

(3.2) 
$$\begin{cases} x^{i} = x^{i}(x^{j}), & \det(\partial x^{i}/\partial \bar{x}^{j}) \neq 0, \\ y^{a} = \frac{\partial x^{a}}{\partial \bar{x}^{b}} \bar{y}^{b}, & i, j, a, b = \overline{1, n}. \end{cases}$$

A generalized gauge transformation on TM is a diffeomorphism of TM compatible with the tangent bundle structure, given locally by

(3.3) 
$$\begin{cases} x^i = X^i(\tilde{x}^j), & \det(\partial X^i/\partial \tilde{x}^j) \neq 0, \\ y^a = Y^a_b(\tilde{x})\tilde{y}^b, & \det(Y^a_b(\tilde{x})) \neq 0. \end{cases}$$

A gauge *d*-tensor field is a *d*-tensor field which obeys tensor-type transformation rules with respect to (3.3); e.g. the gauge *d*-tensor field  $\{w_{jb}^{ia}(x, y)\}$ transforms relative to (3.2) and (3.3) according to

(3.4) 
$$\begin{cases} \bar{w}_{jb}^{i\,a}(\bar{x},\bar{y})\frac{\partial x^{k}}{\partial\bar{x}^{i}}\frac{\partial x^{c}}{\partial\bar{x}^{a}} = w_{l\,d}^{k\,c}(x,y)\frac{\partial x^{l}}{\partial\bar{x}^{j}}\frac{\partial x^{d}}{\partial\bar{x}^{b}},\\ \tilde{w}_{jb}^{ia}(\tilde{x},\tilde{y})\frac{\partial X^{k}}{\partial\bar{x}^{i}}\frac{\partial Y^{c}}{\partial\bar{y}^{a}} = w_{l\,d}^{k\,c}(x,y)\frac{\partial X^{l}}{\partial\bar{x}^{j}}\frac{\partial Y^{d}}{\partial\bar{y}^{b}} \end{cases}$$

Let TM be endowed with the nonlinear connection  $N = \{N_i^a(x, y)\}$ . If  $\delta_k$  given by (1.1) yields *d*-covector fields when acting on gauge scalar fields (i.e. functions  $f \in F(TM)$  which obey  $\bar{f}(\bar{x}, \bar{y}) = f(x, y), \tilde{f}(\tilde{x}, \tilde{y}) =$ f(x, y)), then N is called a generalized gauge non-linear connection. Further, if  $|_k$  and  $|_c$  are the h- and v-covariant derivations associated with a linear *d*-connection D on TM (i.e. a connection that preserves by parallelism the distributions N and VTM [3]), having the coefficients

$$D\Gamma(N) = \left\{ L^{i}_{jk}(x,y), \ L^{a}_{bk}(x,y), \ C^{i}_{ja}(x,y), \ C^{a}_{bc}(x,y) \right\}$$

given by the relations

(3.5) 
$$\begin{cases} D_{\delta_j}\delta_i = L^k_{ij}\delta_k, \quad D_{\dot{\partial}_a}\delta_j = C^i_{ja}\delta_i \\ D_{\delta_k}\dot{\partial}_b = L^a_{bk}\dot{\partial}_a, \quad D_{\dot{\partial}_c}\dot{\partial}_b = C^a_{bc}\dot{\partial}_a \end{cases}$$

then we say that D is a gauge linear d-connection iff  $|_k$  and  $|_c$  preserve the gauge tensorial character.

In the following we suppose that N is given by (A2), that  $g_{ij}$  and  $h_{ab}$ in (3.1) are gauge d-tensor fields, that  $\gamma_{ij}$  and  $\sigma$  in (A1) are gauge fields (tensor, respectively scalar), and we fix the coefficients

$$C_{ja}^{i} = 0, \quad L_{bk}^{a} = \dot{\partial}_{b} N_{k}^{a} = \left\{ \begin{matrix} a \\ bk \end{matrix} \right\}.$$

Then we can infer the following

**Proposition.** If  $g_{ij|k} = 0$  and  $h_{ab}|_c = 0$ ,  $L^i_{jk} = L^i_{kj}$  and  $C^a_{bc} = C^a_{cb}$ , then

(3.6) 
$$\begin{cases} L_{jk}^{i} = \frac{1}{2}g^{is}(\delta_{j}g_{sk} + \delta_{k}g_{sj} - \delta_{s}g_{jk}) \\ C_{bc}^{a} = \frac{1}{2}h^{as}(\dot{\partial}_{b}h_{sc} + \dot{\partial}_{c}h_{sb} - \dot{\partial}_{s}h_{bc}). \end{cases}$$

We remark that if  $g_{ij}$  is given by (A1) and  $h_{ab} = \delta^i_a \delta^j_b \cdot g_{ij}$ , then  $L^i_{jk}$ and  $C_{bc}^{a}$  are those given by (1.1).

It can be shown by direct calculation that the torsion and curvature d-tensor fields of  $D\Gamma(N)$  are gauge d-tensor fields, and we have [3,4]

(3.7) 
$$T^i_{jk} = 0, \quad S^a_{bc} = 0, \quad P^a_{kb} = 0, \quad C^i_{ja} = 0, \quad R^a_{kl} = r^a_{skl} y^s$$

and

(3.8) 
$$\begin{aligned} R^{a}_{bkl} &= r^{a}_{bkl} + C^{a}_{bd} R^{d}_{kl}, \quad P^{i}_{jkc} &= \dot{\partial}_{c} L^{i}_{jk} \\ P^{a}_{bkc} &= -C^{a}_{bc|k}, \quad S^{i}_{jbc} &= 0, \end{aligned}$$

with  $R_{jkl}^i$  and  $S_{bcd}^a$  given by (1.3) for  $L_{jk}^i$  and  $C_{bc}^a$  given by (1.1). So that  $D\Gamma(N)$  admits one torsion and five curvature non-trivial gauge d-tensor fields. It is obvious that the mixed Lagrangian

(3.9) 
$$L = l_1 R + l_5 S + \hat{L}$$

is a gauge scalar field. Here

(3.10) 
$$\hat{L} = l_0 L_0 + l_2 \hat{R} + l_3 \overset{h}{P} + l_4 \overset{v}{P}, \quad L_0 = R^a_{kl} R^{kl}_a, \quad \overset{v}{R} = R^a_{bkl} R^{bkl}_a$$
$$\overset{h}{P} = P^i_{jkc} P^{jkc}_i, \quad \overset{v}{P} = P^a_{bkc} P^{bkc}_a, \quad l_0, l_1, l_2, l_3, l_4, l_5 \in \mathbb{R}$$

and R, S are given by (1.5), the raising/lowering of the indices being performed using  $g_{ij}(x, y)$  and  $h_{ab}(x, y)$  for the corresponding index-types. We remark that the Lagrangian (3.9) depends functionally on the gauge fields  $g_{ii}(x, y)$ ,  $h_{ab}(x, y)$  and their derivatives, by means of (3.8) and (3.6). So that applying the gauge variational principle

$$\delta \int \mathcal{L} dx^n dy^n = 0$$

for the Lagrangian density

(3.11) 
$$\mathcal{L} = LG, \quad G = |\det(g_{ij})|^{1/2} \cdot |\det(h_{ab})|^{1/2},$$

we infer the generalized Einstein-Yang Mills equations for  $\mathcal{L}$ , obtained by vanishing of the Euler-Lagrange derivatives

$$\frac{\delta \mathcal{L}}{\delta \phi} \equiv \partial_k \left( \frac{\partial \mathcal{L}}{\partial (\partial_k \phi)} \right) + \dot{\partial}_a \left( \frac{\partial \mathcal{L}}{\partial (\dot{\partial}_a \phi)} \right) - \frac{\partial \mathcal{L}}{\partial \phi} = 0, \quad \phi \in \{g_{ij}(x, y), h_{ab}(x, y)\}.$$

By direct calculation, we obtain the following equivalent expressions for these equations:

$$R_{ij} - \frac{1}{2}Rg_{ij} = \frac{1}{l_1}g_{il}g_{jm} \left\{ \frac{\partial \hat{L}}{\partial g_{lm}} + g^{st}\frac{\partial R_{st}}{\partial g_{lm}} - \frac{1}{G} \left[ \partial_k \left( G \frac{\partial L}{\partial(\partial_k g_{lm})} \right) + \dot{\partial}_a \left( G \frac{\partial L}{\partial(\dot{\partial}_a g_{lm})} \right) \right] \right\} + \frac{1}{2l_1}g_{ij}(\hat{L} + l_5S),$$
(3.12)

(3.13)

$$S_{ab} - \frac{1}{2}Sh_{ab} = \frac{1}{l_5}h_{ae}h_{bf} \cdot \left\{ \frac{\partial \hat{L}}{\partial h_{ef}} + h^{uv}\frac{\partial S_{uv}}{\partial h_{ef}} - \frac{1}{G} \left[ \partial_k \left( G \frac{\partial L}{\partial(\partial_k h_{ef})} \right) + \dot{\partial}_d \left( G \frac{\partial L}{\partial(\dot{\partial}_d h_{ef})} \right) \right] \right\} + \frac{1}{2l_5}h_{ab}(\hat{L} + l_1R).$$

Remark that, for the case  $h_{ab} = \delta^i_a \delta^j_b \cdot g_{ij}$ , the equations (3.12, 3.13) contain explicitly the Einstein *h*- and *v*-gauge tensor fields considered in (1.6)

(3.14) 
$$\begin{cases} E_{ij} = R_{ij} - \frac{1}{2}Rg_{ij} \\ E_{ab} = S_{ab} - \frac{1}{2}Sh_{ab} \end{cases}$$

and that the right-hand terms of these equations stand for the energymomentum h- and v-tensor fields, respectively.

In order to obtain solutions  $\{g_{ij}(x, y), h_{ab}(x, y)\}$  for (3.12) and (3.13) having the predefined form (A1), a necessary condition will be (as L will depend on  $\sigma(x, y)$  and its derivatives) the vanishing of the corresponding Euler-Lagrange derivative

(3.15) 
$$\frac{\delta L}{\delta \sigma} \equiv \frac{1}{G} \left\{ \partial_k \left[ \frac{\partial \mathcal{L}}{\partial (\partial_k \sigma)} \right] + \dot{\partial}_a \left( \frac{\partial \mathcal{L}}{\partial \dot{\sigma}_a} \right) - \frac{\partial \mathcal{L}}{\partial \sigma} \right\} = 0.$$

We shall describe the generalised Einstein-Yang Mills equation (3.15) for the gauge Lagrangian L given in (3.9) for three special cases. 1°. If  $g_{ij}(x,y) = e^{2\sigma(x,y)}\gamma_{ij}(x)$  and  $h_{ab} = \delta^i_a \delta^j_b g_{ij}$ , then we obtain the case of the almost Hermitian model  $H^{2n} = (TM, G_*, F)$  [3] given by the

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N-lift of the generalized Lagrange metric (A1) to TM, and the almost complex structure F on TM given in the local adapted frame (1.1") by

$$F(\delta_i) = -\dot{\partial}_i, \quad F(\dot{\partial}_i) = \delta_i.$$

In this case, we remark that  $L_{jk}^i$  and  $C_{bc}^a$  are given by (1.1), R and S are described in (1.5), and (3.15) becomes

$$(3.16) \quad \frac{\delta L}{\delta \sigma} = \frac{1}{G} \left[ \partial_k \left( l_1 H \frac{\partial \sigma}{\partial \beta_k} - 2l_4 G \frac{\partial (C_{bc|l}^a)}{\partial \beta_k} P_a^{blc} \right) + \dot{\partial}_a \left( G \frac{\partial \bar{L}}{\partial \dot{\sigma}_a} \right) \right] + 2(1-n)\bar{L} + 2(\bar{L} - l_5 S) = 0$$

where  $H = Ge^{-2\sigma}$ ,  $\beta_k = \partial_k \sigma$ , and

(3.17) 
$$\bar{L} = l_2 \overset{v}{R} + l_3 \overset{h}{P} + l_4 \overset{v}{P} + l_5 S.$$

2°. For  $g_{ij} = \gamma_{ij}(x)$ ,  $h_{ab} = e^{2\sigma(x,y)}\gamma_{ab}(x)$ , we have  $L^i_{jk} = {i \\ jk}$ , R = r,  $C^a_{bc}$  given by (1.1-2), and

(3.18) 
$$\frac{\delta L}{\delta \sigma} = \frac{1}{G} \left[ -2l_4 \partial_k \left( GP_a^{blc} \frac{\partial (C_{bc|l}^a)}{\partial \beta_k} \right) + \dot{\partial}_a \left( G \frac{\partial \bar{L}}{\partial \dot{\sigma}_a} \right) \right] + 2(l_4 \overset{v}{P} + l_5 S - l_0 L_0) - nL = 0.$$

3°. For  $g_{ij}(x,y) = e^{2\sigma(x,y)}\gamma_{ij}(x)$ ,  $h_{ab} = \gamma_{ab}(x)$ , we have  $L^i_{jk}$  given by (1.1-1), P = S = 0, and

$$(3.19) \qquad \qquad \frac{\delta L}{\delta \sigma} = \frac{1}{G} \left\{ l_1 \partial_k \left( H \frac{\partial \sigma}{\partial \beta_k} \right) + 2 \left[ l_2 \frac{\partial}{\partial y^a} \left( G R_f^{bkl} R_{kl}^d (\delta_b^f \delta_d^a + \delta_d^f \delta_b^a - \gamma_{bd} \gamma^{af}) \right) + l_3 \frac{\partial}{\partial y^a} \left( G P_i^{jkc} (\gamma_{kj} \gamma^{is} N_s^a - \delta_j^i N_k^a - \delta_k^i N_j^a) \right) \right] \right\} + 2(2l_0 L_0 + l_1 R + 2l_2 \overset{v}{R} + l_3 \overset{h}{P}) - n\bar{L} = 0.$$

An open problem is the one of the one of determining valid solutions  $\sigma(x, y)$  for the equation (3.15), for suitable constants  $l_0, \ldots, l_5$  in the Lagrangian field (3.9).

Conclusions. The space  $GL^n$ , as a model for the geometry of spacetime [6] can be examined from the point of view of determining its Einstein and Maxwell equations, using for the first ones two approaches: the theory of *d*-object fields for generalised Lagrange spaces and the generalized gauge theory for vector bundles endowed with (h, v)-metrics. Explicit forms for these equations are obtained, as a preliminary step for determining their solutions. The Einstein-Yang Mills equation with respect to the scalar gauge field  $\sigma(x, y)$  is also derived for three cases of (h, v)-metrics related to the fundamental metric of  $GL^n$ .

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