

## On a class of Einstein space-time manifolds

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**Abstract.** We deal with a general space-time  $(M, g)$  with usual differentiability conditions and hyperbolic metric  $g$  of index 1, which carries 3 skew-symmetric Killing vector fields  $X, Y, Z$  having as generative the unit time-like vector field  $e$  of the hyperbolic metric  $g$ . It is shown that such a space-time  $(M, g)$  is an Einstein manifold of curvature  $-1$ , which is foliated by space-like hypersurfaces  $M_s$  normal to  $e$  and the immersion  $x : M_s \rightarrow M$  is pseudo-umbilical. In addition, it is proved that the vector fields  $X, Y, Z$  and  $e$  are exterior concurrent vector fields and  $X, Y, Z$  define a commutative Killing triple,  $M$  admits a Lorentzian transformation which is in an orthochronous Lorentz group and the distinguished spatial 3-form of  $M$  is a relatively integral invariant of the vector fields  $X, Y$  and  $Z$ .

### 0. Introduction

Let  $(M, g)$  be a general space-time with usual differentiability conditions and hyperbolic metric  $g$  of index 1.

We assume in this paper that  $(M, g)$  carries 3 skew-symmetric Killing vector fields (abbr. SSK)  $X, Y, Z$  having as generative the unit time-like vector field  $e$  of the hyperbolic metric  $g$  (see [R1], [MRV]). Therefore, if  $\nabla$  is the Levi-Civita connection and  $\wedge$  means the wedge product of vector

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fields, one has

$$\nabla X = X \wedge e, \quad \nabla Y = Y \wedge e, \quad \nabla Z = Z \wedge e. \quad (0.1)$$

Setting  $\alpha = X^\flat$ ,  $\beta = Y^\flat$ ,  $\gamma = Z^\flat$  ( $\flat : TM \rightarrow T^*M$ ,  $X \mapsto X^\flat$ , denotes the musical isomorphism defined by  $g$ ), one derives by (0.1)

$$d\alpha = 2\omega \wedge \alpha, \quad \omega = e^\flat, \quad (0.2)$$

and clearly similar equations for  $\beta$  and  $\gamma$ .

Since one finds that  $\omega$  is an exact form, equations (0.2) may be written in terms of  $d^\omega$ -cohomology as  $d^{-2\omega}\alpha = 0$  and say that  $\alpha$ ,  $\beta$ ,  $\gamma$  are  $d^{-2\omega}$ -exact forms. From (0.2) it also follows that the existence of such a space-time is determined by an exterior differential system in involution [C2].

The following theorem is proved:

Any space-time  $(M, g)$  satisfying (0.1) is an Einstein manifold of curvature  $-1$ , which is foliated by space-like hypersurfaces  $M_s$  normal to  $e$  and the immersion  $x : M_s \rightarrow M$  is pseudo-umbilical [Ch].

The following additional properties are also obtained.

- i) the vector fields  $X$ ,  $Y$ ,  $Z$  and  $e$  are exterior concurrent vector fields (abbr. EC) and  $X$ ,  $Y$ ,  $Z$  define a commutative Killing triple;
- ii)  $M$  admits a Lorentzian transformation which is in an orthochronous Lorentz group [CWD];
- iii) the 3-form  $\varphi = \alpha \wedge \beta \wedge \gamma$  is a relatively integral invariant of the vector fields  $X$ ,  $Y$  and  $Z$ .

## 1. Preliminaries

Let  $(M, g)$  be a pseudo-Riemannian  $C^\infty$ -manifold and let  $\nabla$  be the covariant differential operator defined by the metric tensor  $g$ . We assume that  $M$  is oriented and that  $\nabla$  is the Levi-Civita connection. Let  $\Gamma TM$  be the set of sections of the tangent bundle  $TM$  and  $\flat : TM \rightarrow T^*M$  and  $\sharp : T^*M \rightarrow TM$  the classical musical isomorphisms defined by  $g$ .

Following [P], we set  $A^q(M, TM) = \Gamma \text{Hom}(\Lambda^q TM, TM)$  and notice that elements of  $A^q(M, TM)$  are vector valued  $q$ -forms,  $q \leq \dim M$ .

Denote by  $d^\nabla : A^q(M, TM) \rightarrow A^{q+1}(M, TM)$  the *exterior covariant derivative operator* with respect to  $\nabla$  (it should be noticed that generally  $d^{\nabla^2} = d^\nabla \circ d^\nabla \neq 0$ , unlike  $d^2 = d \circ d = 0$ ).

If  $p \in M$ , then the vector valued 1-form  $dp \in A^1(M, TM)$  is called *the soldering form* of  $M$  ( $dp$  is the canonical vector valued 1-form of  $M$  and one has  $d^\nabla(dp) = 0$ ). The operator  $d^\omega = d + e(\omega)$  acting on  $\Lambda M$  is called *the cohomology operator*, where  $e(\omega)$  means the exterior product by the closed 1-form  $\omega \in \Lambda^1 M$ , i.e.  $d^\omega u = du + \omega \wedge u$ , for any  $u \in \Lambda M$ . We have  $d^\omega \circ d^\omega = 0$  and if  $d^\omega u = 0$ ,  $u$  is said to be  *$d^\omega$ -closed* [GL].

Any vector field  $X \in \Gamma TM$  such that

$$d^\nabla(\nabla X) = \nabla^2 X = \pi \wedge dp \in A^2(M, TM)$$

is said to be an *exterior concurrent* vector field.

The 1-form  $\pi$ , which is called the *concurrence form*, is given by

$$\pi = f X^\flat, \quad f \in C^\infty M.$$

If  $\mathcal{R}$  denotes the *Ricci tensor* of  $\nabla$ , we have

$$\mathcal{R}(X, Y) = -(n - 1)fg(X, Z), \quad Z \in \Gamma TM, \quad n = \dim M,$$

and consequently

$$f = -\frac{1}{n - 1} \text{Ric}(X),$$

where  $\text{Ric}(X)$  means the *Ricci curvature* of  $M$  with respect to  $X$ .

Let  $\mathcal{O} = \{e_A \mid A \in \{1, \dots, n\}\}$  be an adapted local field of orthonormal frames on  $M$  and let  $\mathcal{O}^* = \{\omega^A\}$  be its associated coframe.

With respect to  $\mathcal{O}$  and  $\mathcal{O}^*$ , the soldering form  $dp$  and E. Cartan's structure equations in indexless form are

$$dp = \omega^A \otimes e_A \in A^1(M, TM), \tag{1.1}$$

$$\nabla e = \theta \otimes e \in A^1(M, TM), \tag{1.2}$$

$$d\omega = -\theta \wedge \omega, \tag{1.3}$$

$$d\theta = -\theta \wedge \theta + \Theta. \tag{1.4}$$

In the above equations,  $\theta$ , respectively  $\Theta$ , are the *local connection forms* in the bundle  $\mathcal{O}(M)$ , respectively the *curvature forms* on  $M$ .

## 2. Main result

Let  $(M, g)$  be a general space-time with usual differentiability conditions and normal *hyperbolic* metric  $g$ , i.e. of index 1. Let  $\mathcal{O} = \{e_A \mid A \in \{1, \dots, n\}\}$  be an adapted local field of orthonormal frames on  $M$  and let  $\mathcal{O}^* = \{\omega^A\}$  be its associated coframe. We agree to denote by  $e_a$ ,  $a, b \in \{1, 2, 3\}$  and by  $e_4$  the space-like vector basis and the time-like vector basis, respectively, w.r.t.  $g$ . Then, by reference to [C1] (see also [MRV]), one has

$$dp = -\omega^a \otimes e_a + \omega^4 \otimes e_4 \implies \langle dp, dp \rangle = (\omega^4)^2 - \sum_{a=1}^3 (\omega^a)^2, \quad (2.1)$$

and Cartan's structure equations are expressed by

$$\begin{cases} \nabla e_a = -\theta_a^b e_b + \theta_a^4 \otimes e_4, \\ \nabla e_4 = -\theta_4^a \otimes e_a, \end{cases} \quad (2.2)$$

$$\begin{cases} d\omega^a = -\omega^b \wedge \theta_b^a + \omega^4 \otimes \theta_4^a, \\ d\omega^4 = -\omega^a \wedge \theta_a^4, \end{cases} \quad (2.3)$$

$$\begin{cases} d\theta_a^b = \Theta_a^b - \theta_a^c \wedge \theta_c^b + \theta_a^4 \wedge \theta_4^b, \\ d\theta_4^a = \Theta_4^a - \theta_4^c \wedge \theta_c^a. \end{cases} \quad (2.4)$$

In the following, in order to simplify, we set  $\omega^4 = \omega$  and  $e_4 = e$ .

In this paper we assume that the manifold  $M$  under consideration carries 3 space-like vector fields  $X, Y, Z$  which are skew-symmetric Killing vector fields (abbr. SSK) [R1], [MRV] having as generative the unit time-like vector field  $e$ . In order to simplify, we also set

$$X^b = \alpha, \quad Y^b = \beta, \quad Z^b = \gamma. \quad (2.5)$$

Under these conditions, by reference to [R1], [MRV], one has

$$\begin{cases} \nabla X = X \wedge e = \omega \otimes X - \alpha \otimes e, \\ \nabla Y = Y \wedge e = \omega \otimes Y - \beta \otimes e, \\ \nabla Z = Z \wedge e = \omega \otimes Z - \gamma \otimes e, \end{cases} \quad (2.6)$$

and since  $X, Y, Z$  are space-like, we set

$$X = X^a e_a, \quad Y = Y^a e_a, \quad Z = Z^a e_a, \quad (2.7)$$

i.e.

$$\|X\|^2 = -\sum_{a=1}^3 (X^a)^2, \quad \|Y\|^2 = -\sum_{a=1}^3 (Y^a)^2, \quad \|Z\|^2 = -\sum_{a=1}^3 (Z^a)^2. \quad (2.8)$$

From (2.2) and (2.6), one derives

$$dX^a - X^b \theta_b^a = X^a \omega \quad (2.9)$$

and

$$\alpha = -X^a \theta_a^4. \quad (2.10)$$

Hence, it follows from (2.10)

$$\theta_a^4 = -\omega^a \quad (2.11)$$

Setting  $2\varphi_x = \sum_{a=1}^3 (X^a)^2$ ,  $2\varphi_y = \sum_{a=1}^3 (Y^a)^2$ ,  $2\varphi_z = \sum_{a=1}^3 (Z^a)^2$ , one gets at once from (2.9)

$$\frac{d\varphi_x}{2\varphi_x} = \omega, \quad \frac{d\varphi_y}{2\varphi_y} = \omega, \quad \frac{d\varphi_z}{2\varphi_z} = \omega, \quad (2.12)$$

which shows that the generative  $\omega$  (i.e. the time-like covector) is exact. Further, taking the exterior differential of  $\alpha$ , one finds by (2.9) and by the structure equations (2.3)

$$d\alpha = 2\omega \wedge \alpha, \quad (2.13)$$

which is the general equation of a SSK vector field [R1].

Now, by (2.11) and the second equation (2.4), one derives on behalf of the structure equations (2.3) regarding the space-like covectors  $\omega^a$

$$\Theta_4^a = -\omega^a \wedge \omega. \quad (2.14)$$

Further, taking the exterior differentials of (2.4)

$$\begin{cases} X^2(\Theta_1^2 - \omega^1 \wedge \omega^2) + X^3(\Theta_1^3 - \omega^1 \wedge \omega^3) = 0, \\ X^1(\Theta_2^1 - \omega^2 \wedge \omega^1) + X^3(\Theta_2^3 - \omega^2 \wedge \omega^3) = 0. \end{cases}$$

Since, clearly, similar equations hold for  $Y$  and  $Z$ , one infers

$$\Theta_b^a = \omega^b \wedge \omega^a. \quad (2.15)$$

So, by reference to a known formula regarding space forms, we conclude by (2.14) and (2.15) that the space-time manifold  $(M, g)$  under consideration is an *Einstein manifold* of curvature  $-1$ .

Since  $d\omega = 0$ , it is seen that  $(M, g)$  is foliated by space-like hypersurfaces  $M_s$  and by (2.11) and the equation (2.2), one has

$$\nabla e = \omega^a \otimes e_a = -dp_s, \quad (2.16)$$

where  $dp_s$  is the *spatial component* of the soldering form  $dp$ . Then the second fundamental form  $II = -\langle dp_s, dp_s \rangle$  associated with the immersion  $x : M_s \rightarrow M$  being conformal to the metric tensor  $g_s$  of  $M_s$ , it follows [Ch] that the immersion  $x$  is *pseudo-umbilical*.

On the other hand, clearly,  $Y$  and  $Z$  enjoy the same properties as  $X$  and one may write

$$\begin{cases} d\beta = 2\omega \wedge \beta \Leftrightarrow d^{-2\omega}\beta = 0, \\ d\gamma = 2\omega \wedge \gamma \Leftrightarrow d^{-2\omega}\gamma = 0 \end{cases} \quad (2.17)$$

(since  $\omega$  is an exact form one may say that, cohomologically, the dual forms  $\alpha, \beta, \gamma$  of the SSK vector fields  $X, Y, Z$  are  $d^{-2\omega}$ -exact).

Finally, by (2.12), (2.13) and (2.17), it is seen that the existence of the considered space-time is defined by an exterior differential system  $\Sigma$ , whose characteristic numbers [C2] are  $r = 4, s_0 = 1, s_1 = 3$ . Consequently, we conclude that  $\Sigma$  is in *involution*, in the sense of CARTAN [C2], and depends of 3 arbitrary functions of one argument.

Since the vector fields  $X, Y, Z$  are orthogonal to  $e$ , one easily finds by (2.6)

$$[X, Y] = 0, \quad [X, Z] = 0, \quad [Y, Z] = 0,$$

which proves that the SSK vector fields  $X, Y, Z$  define a *commutative triple*.

Summing up, we state the following.

**Theorem.** *Let  $(M, g)$  be a space-time manifold with normal hyperbolic metric  $g$ . Assume that  $M$  carries 3 skew-symmetric Killing vector*

fields  $X, Y, Z$  having as generative the unit time-like vector field  $e$  of the hyperbolic metric  $g$ . Then  $M$  is an Einstein manifold of curvature  $-1$ . The existence of the triple  $\{X, Y, Z\}$  is assured by an exterior differential system  $\Sigma$  in involution. Such a manifold  $(M, g)$  has also the following properties:

- i)  $M$  is foliated by space-like hypersurfaces  $M_s$  tangent to  $\{X, Y, Z\}$ , normal to  $e$  and the immersion  $x : M_s \rightarrow M$  is pseudo-umbilical;
- ii) the vector fields  $\{X, Y, Z\}$  define a commutative triple of Killing vector fields.

### 3. Additional properties

In this section we shall make some additional considerations regarding the Einstein manifold defined in Section 2.

By using (2.16) and operating by  $d^\nabla$ , one may write

$$d^\nabla(dp_s) = \omega \wedge dp_s. \tag{3.1}$$

Because  $dp = \omega \otimes e - dp_s$ , one derives

$$\nabla^2 e = \omega \wedge dp. \tag{3.2}$$

By reference to [R1], [PRV], the above equation proves that  $e$  is an EC vector field. Hence, following the general theory [PRV], if  $W$  is any vector field on  $M$ , one may write  $\mathcal{R}(e, W) = -3g(e, W)$  (we notice that for any space-like vector field  $Z_s$  one has  $\mathcal{R}(e, Z_s) = 0$ ).

Recall that the sectional curvature  $K_{U \wedge V}$  of any vector fields  $U, V$  is expressed by

$$K_{U \wedge V} = \frac{g(R(U, V)V, U)}{\|U\|^2 \|V\|^2 - g(U, V)^2}. \tag{3.3}$$

Then by (2.6) and recalling that  $X, Y, Z$  are space-like vector fields, one finds  $K_{X \wedge Y} = K_{Y \wedge Z} = K_{Z \wedge X} = -1$ , which means that  $M$  is of curvature  $-1$ .

Next, operating by  $d^\nabla$  on the vector fields  $X, Y, Z$ , one derives

$$\nabla^2 X = \alpha \wedge dp, \quad \nabla^2 Y = \beta \wedge dp, \quad \nabla^2 Z = \gamma \wedge dp. \tag{3.4}$$

From (3.4) it follows that the SSK vector fields  $X$ ,  $Y$ ,  $Z$  are also exterior concurrent, as the time-like vector field  $e$ .

Next, setting  $\alpha = \lambda U^1$ ,  $\beta = \lambda U^2$ ,  $\gamma = \lambda U^3$  as a space-like covector basis, one may define a subgroup of Lorentz by the group of space-like rotations  $O(3)$  (*orthochronous transformations* [CWD]) preserving the time-like vector field  $e$ .

Also, one has

$$U^4 = \omega^4 = \omega, \quad \alpha^2 + \beta^2 + \gamma^2 = \lambda^2 \sum_{a=1}^3 (\omega^a)^2, \quad (3.5)$$

where  $\lambda$  is a scalar field.

Equations (3.5) imply

$$\sum_{a=1}^3 [(X^a)^2 + (Y^a)^2 + (Z^a)^2] = \lambda^2 \quad (3.6)$$

and

$$\left\{ \begin{array}{l} \sum_{a \neq b} [X^a X^b + Y^a Y^b + Z^a Z^b] = 0, \\ \sum_{a \neq b} [X^a X^c + Y^a Y^c + Z^a Z^c] = 0, \\ \sum_{a \neq b} [X^b X^c + Y^b Y^c + Z^b Z^c] = 0. \end{array} \right. \quad (3.7)$$

Making use of equations (2.9), one finds that the differentiation of (3.7) holds good and the differentiation of (3.6) gives

$$\frac{d\lambda}{\lambda} = \omega \quad (3.8)$$

(recall that  $\omega$  is an exact form).

Finally, we shall outline a certain property of the Lie algebra induced by the space-like vector fields  $X$ ,  $Y$ ,  $Z$ . We agree to call the 3-form

$$\varphi = \alpha \wedge \beta \wedge \gamma \quad (3.9)$$

the *distinguished spatial form* of  $M$ .



By (2.13) one gets at once

$$d\varphi = 6\omega \wedge \varphi \iff d^{-6\omega}\varphi = 0 \iff \mathcal{L}_e\varphi = 6\varphi, \tag{3.10}$$

which shows that  $\varphi$  is an  $d^{-6\omega}$ -exact form and  $e$  defines an *infinitesimal conformal transformation* of  $\varphi$ .

By (2.12) one has

$$\frac{dg(X, X)}{4g(X, X)} = \frac{dg(Y, Y)}{4g(Y, Y)} = \frac{dg(Z, Z)}{4g(Z, Z)} = \omega, \tag{3.11}$$

and one derives

$$i_X\varphi = g(X, X)\beta \wedge \gamma + g(X, Y)\gamma \wedge \alpha + g(X, Z)\alpha \wedge \beta$$

and similar relations for  $Y$  and  $Z$ .

So, by (2.13), one infers  $d(i_X\varphi) = 8\omega \wedge i_X\varphi$ , which gives  $\mathcal{L}_X\varphi = 2\omega \wedge i_X\varphi$ .

Finally, by exterior differentiation, one may write  $d(\mathcal{L}_X\varphi) = 0$  and clearly similar relations for  $Y$  and  $Z$  hold.

Hence, following a known definition [A], one may state that the distinguished spatial 3-form  $\varphi$  is a *relatively integral invariant* of the SSK vector fields  $X, Y, Z$ .

Consequently, the following results were obtained.

**Theorem.** *Let  $(M, g)$  be the space-time manifold defined in the Section 2 and let  $X, Y, Z$  be the 3 skew-symmetric Killing vector fields which determine  $M$  and  $e$  the unit time-like vector field of the hyperbolic metric  $g$ . One has the following properties:*

- i) *the vector fields  $X, Y, Z$  and  $e$  are exterior concurrent vector fields;*
- ii)  *$M$  admits an orthogonal transformation of a space-like Lorentz subgroup  $O(3)$ ;*
- iii) *the 3-form  $\varphi = \alpha \wedge \beta \wedge \gamma$  is a relatively integral invariant of the vector fields  $X, Y$  and  $Z$ .*

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