

## Spectral geometry on certain almost Hermitian Einstein manifolds

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**Abstract.** On compact Riemannian and Kähler manifolds the spectra of the real and complex Laplacians determine the geometry of the manifolds to a considerable extent, though not completely, as isospectral manifolds need not be isometric. The literature on the geometric consequences of isospectrality is extensive (e.g., [1], [2], [4]–[8], [10], [11]). In this paper we consider the spectrum of the real Laplacian on three classes of almost Hermitian Einstein manifolds which include almost Kähler and nearly Kähler Einstein manifolds. We prove that a compact almost Hermitian Einstein manifold of positive scalar curvature  $\varrho$  with  $\varrho \leq \varrho^*$ , or negative scalar curvature  $\varrho$  with  $\varrho \geq \varrho^*$ , or nonzero scalar curvature  $\varrho$  with  $\varrho = \varrho^*$  (where  $\varrho^*$  is the  $*$ -scalar curvature) and isospectral to a Kähler manifold of constant holomorphic sectional curvature is Kähler and the manifolds are holomorphically isometric.

### 1. Preliminaries

Let  $(M, g)$  be a Riemannian manifold of real dimension  $m = 2n \geq 2$  with metric  $g = (g_{ij})$ . If  $R = (R_{hijk})$  is the Riemann curvature tensor,  $Rc = (R_{ij}) = g^{hk} R_{hijk}$  the Ricci curvature tensor and  $\varrho = g^{ij} R_{ij}$  the scalar curvature, then the Einstein tensor  $E = (E_{ij})$  is given by

$$(1.1) \quad E_{ij} \equiv R_{ij} - \frac{\varrho}{m} g_{ij},$$

where  $M$  is Einstein if  $E = 0$ .

If  $(M, g)$  is a compact connected  $C^\infty$  manifold and  $\Delta = -(d\delta + \delta d)$  is the Laplace operator on  $p$ -forms,  $0 \leq p \leq 2n$ , (0-forms corresponding to differentiable functions on  $M$ ) with respect to the metric  $g$ , then the spectrum of the Laplacian are the eigenvalues of  $\Delta$ ,

$$(1.2) \quad \text{Spec}^p(M, g) = \{\lambda_{ip} \mid 0 \geq \lambda_{1,p} \geq \lambda_{2,p} \geq \cdots \geq \lambda_{k,p} \geq \cdots \downarrow -\infty\},$$

where each eigenvalue is repeated as often as its multiplicity. Further,  $\text{Spec}^{2n-p}(M, g) = \text{Spec}^p(M, g)$  when  $M$  is orientable.

Relevant to the study of the spectrum is the Minakshisundaram-Pleijel-Gaffney asymptotic formula

$$(1.3) \quad \sum_{k=0}^{\infty} \exp(\lambda_{k,p} t) \underset{t \rightarrow 0}{\sim} \frac{1}{(4\pi t)^n} \sum_{i=0}^{\infty} a_{i,p} t^i,$$

where the first three coefficients are given by [8]

$$(1.4) \quad a_{0,p} = \binom{2n}{p} \int_M dM = \binom{2n}{p} \text{vol}(M),$$

$$(1.5) \quad a_{1,p} = \left[ \frac{1}{6} \binom{2n}{p} - \binom{2n-2}{p-1} \right] \int_M \varrho dM,$$

$$(1.6) \quad a_{2,p} = \int_M [c_1(2n, p) \varrho^2 + c_2(2n, p) |Rc|^2 + c_3(2n, p) |R|^2] dM;$$

where

$$(1.7) \quad c_1(2n, p) = \frac{1}{72} \binom{2n}{p} - \frac{1}{6} \binom{2n-2}{p-1} + \frac{1}{2} \binom{2n-4}{p-2},$$

$$(1.8) \quad c_2(2n, p) = -\frac{1}{180} \binom{2n}{p} + \frac{1}{2} \binom{2n-2}{p-1} - 2 \binom{2n-4}{p-2},$$

$$(1.9) \quad c_3(2n, p) = \frac{1}{180} \binom{2n}{p} - \frac{1}{12} \binom{2n-2}{p-1} + \frac{1}{2} \binom{2n-4}{p-2}.$$

## 2. Almost Hermitian manifolds and the Bochner curvature tensor

Let  $(M, g, J)$  be an almost Hermitian manifold of real dimension  $2n \geq 2$  with almost complex structure  $J = (F_i^j)$  and almost Hermitian metric  $g = (g_{ij})$ ; that is,  $g(JX, JY) = g(X, Y)$  for all  $X, Y$  in the tangent space of  $M$  at  $p$ ,  $T_p(M)$ . In dimension  $2n = 2$ ,  $(M, g)$  is Kähler Einstein and has holomorphic sectional curvature  $\kappa = \frac{\varrho}{2}$ .

Define the Bochner curvature tensor  $B = (B_{hijk})$  by

$$(2.1) \quad \begin{aligned} B_{hijk} \equiv & R_{hijk} - \frac{1}{2n+4} (R_{ij}g_{hk} - R_{ik}g_{hj} + R_{hk}g_{ij} - R_{hj}g_{ik} \\ & + F_{ij}F_h^r R_{rk} - F_{ik}F_h^r R_{rj} + F_{hk}F_i^r R_{rj} - F_{hj}F_i^r R_{rk} \\ & - 2F_{jk}F_h^r R_{ri} - 2F_{hi}F_j^r R_{rk}) \end{aligned}$$

$$+ \frac{\varrho}{(2n+2)(2n+4)} (g_{ij}g_{hk} - g_{ik}g_{hj} + F_{ij}F_{hk} - F_{ik}F_{hj} - 2F_{hi}F_{jk}) .$$

**Lemma 2.1.** (see, e.g., [3]). *If  $(M, g)$  is a Kähler manifold of nonzero constant holomorphic sectional curvature  $\kappa$ , then the Riemann curvature tensor, Ricci curvature tensor and scalar curvature are given, respectively, by*

$$(2.2) \quad R_{hijk} = \frac{\kappa}{4} (g_{hk}g_{ij} - g_{hj}g_{ik} + F_{hk}F_{ij} - F_{hj}F_{ik} - 2F_{hi}F_{jk}) ,$$

$$(2.3) \quad R_{ij} = \frac{n+1}{2} \kappa g_{ij} ,$$

$$(2.4) \quad \varrho = n(n+1)\kappa .$$

Hence,  $(M, g)$  is Einstein and  $B = 0$ .

By Schur's theorem, a Kähler manifold of dimension  $2n \geq 4$  with constant holomorphic sectional curvature  $\kappa$  is of constant holomorphic curvature; that is,  $\kappa$  is a global constant on the manifold.

**Lemma 2.2.** [3]. *An almost Hermitian manifold  $(M, g)$  of dimension  $2n \geq 4$  with curvature tensor given by (2.2) is Kähler and of constant holomorphic sectional curvature  $\kappa = \frac{\varrho}{n(n+1)}$ .*

Consequently, we have

**Corollary 2.3.** *If  $(M, g)$  is an almost Hermitian Einstein manifold of dimension  $2n \geq 4$  with  $B = 0$  and  $\varrho \neq 0$ , then the conclusion of Lemma 2.2 follows.*

**Lemma 2.4.** *If  $(M, g)$  is an almost Hermitian Einstein manifold, then the square length of the Bochner tensor is given by*

$$(2.5) \quad |B|^2 = |R|^2 + \frac{\varrho^2 - 3\varrho\varrho^*}{n(n+1)} ,$$

where  $\varrho^* \equiv F^{kq}F^{jr}R_{krqj}$ .

PROOF. The equation follows from a calculation of  $|B_{hijk}|^2$  using (1.1) and (2.1).

### 3. Spectral geometry on almost Hermitian Einstein manifolds

An almost Kähler manifold  $(M, g, J)$  is an almost Hermitian manifold such that the differential form  $\omega \equiv F_{ij}dx^i \wedge dx^j$  is closed; that is,  $d\omega = 0$

and hence,  $\delta\omega = 0$  and  $\omega$  is harmonic. This is equivalent to  $\nabla_i F_{jk} + \nabla_j F_{ki} + \nabla_k F_{ij} = 0$ , where  $\nabla$  is the covariant derivative with respect to the Riemannian connection on  $M$ . A nearly Kähler manifold is an almost Hermitian manifold satisfying  $\nabla_i F_j^k + \nabla_j F_i^k = 0$ .

**Lemma 3.1.** (see, e.g., [3]). *On an almost Kähler manifold*

$$(3.1) \quad \varrho \leq \varrho^* ;$$

*while on a nearly Kähler manifold*

$$(3.2) \quad \varrho \geq \varrho^* .$$

*Further, equality holds in each case if and only if the manifold is Kähler.*

**Theorem 3.2.** *Suppose that  $(M, g, J)$  and  $(M', g', J')$  are compact almost Hermitian manifolds with  $\text{Spec}^p(M, g) = \text{Spec}^p(M', g')$  for  $p = 0, 1$  or  $2$ .*

a) *In dimension  $2n = 2$ ,  $(M, g)$  is of constant holomorphic curvature  $\kappa$  if and only if  $(M', g')$  is.*

b) *In dimension  $2n \geq 4$ , if  $(M', g')$  is Einstein and of positive scalar curvature  $\varrho'$  with  $\varrho' \leq \varrho'^*$ , or negative scalar curvature  $\varrho'$  with  $\varrho' \geq \varrho'^*$ , or nonzero scalar curvature  $\varrho'$  with  $\varrho' = \varrho'^*$ , and  $(M, g)$  is Kähler and of constant holomorphic sectional curvature  $\kappa$ , then  $(M', g')$  is a Kähler manifold of constant holomorphic sectional curvature  $\kappa' = \kappa$  in the following cases:  $p = 0$  and  $2n \geq 4$ ;  $p = 1$  and  $2n \geq 16$ ;  $p = 2$  and  $2n = 4, 6, 8, 14$  or  $2n \geq 18$ .*

PROOF. Letting  $p = 0$  in (1.4)–(1.9) gives

$$(3.3) \quad a_{0,0} = \int_M dM = \text{vol}(M) ,$$

$$(3.4) \quad a_{1,0} = \frac{1}{6} \int_M \varrho dM ,$$

$$(3.5) \quad a_{2,0} = \frac{1}{360} \int_M [5\varrho^2 - 2|Rc|^2 + 2|R|^2] dM .$$

a) In dimension  $2n = 2$ ,  $|R|^2 = \varrho^2$  and since  $g$  is an Einstein metric, then by (1.1),  $|Rc|^2 = \frac{\varrho^2}{2}$  and  $a_{2,0} = \frac{1}{60} \int_M \varrho^2 dM$ . Since  $a_{0,0} = a'_{0,0}$ , it follows that  $\text{vol}(M) = \text{vol}(M')$ . If, say,  $(M, g)$  has constant holomorphic curvature  $\kappa$ , since  $a_{1,0} = a'_{1,0}$  and  $a_{2,0} = a'_{2,0}$ , then  $\int_{M'} \varrho' dM' = 2\kappa \text{vol}(M')$  and  $\int_{M'} \varrho'^2 dM' = 4\kappa^2 \text{vol}(M')$ . We then have equality in the Schwarz inequality  $(\int_{M'} \varrho' dM')^2 \leq (\int_{M'} \varrho'^2 dM') (\int_{M'} dM')$ , so that  $\varrho'$  is constant and  $\varrho' = \varrho$ .

The proofs in the remaining cases are similar upon taking  $p = 1$  and  $2$ , respectively, in (1.4)–(1.9). For  $p = 2$  the result follows, as well, as a consequence of the case  $p = 0$  since  $\text{Spec}^0(M, g) = \text{Spec}^0(M', g')$  by duality.

b) Suppose  $(M', g')$  is almost Hermitian Einstein of either positive scalar curvature  $\varrho'$  with  $\varrho' \leq \varrho'^*$  or negative scalar curvature  $\varrho'$  with  $\varrho' \geq \varrho'^*$ . In dimension  $2n \geq 4$ , substitute (2.5) in (3.5). Then, since  $E = 0$  and  $a_{2,0} = a'_{2,0}$ , we have

$$\begin{aligned}
& \int_M \left[ \frac{(5n^2 + 4n - 3)\varrho^2 + 6\varrho\varrho^*}{n(n+1)} + 2|B|^2 \right] dM \\
&= \int_M \left[ \frac{(5n^2 + 4n + 3)\varrho^2}{n(n+1)} + 2|B|^2 \right] dM \\
(3.6) \quad &= \int_{M'} \left[ \frac{(5n^2 + 4n - 3)\varrho'^2 + 6\varrho'\varrho'^*}{n(n+1)} + 2|B'|^2 \right] dM' \\
&\geq \int_{M'} \left[ \frac{(5n^2 + 4n + 3)\varrho'^2}{n(n+1)} + 2|B'|^2 \right] dM'.
\end{aligned}$$

Since  $(M, g, J)$  is Kähler with constant holomorphic sectional curvature  $\kappa$ , then by Lemma 2.1,  $B = 0$ . Further, since  $\varrho$  is constant, then  $a_{0,0} = a'_{0,0}$ ,  $a_{1,0} = a'_{1,0}$  and the Schwarz inequality imply that  $\left( \int_{M'} \varrho'^2 dM' \right) \left( \int_{M'} dM' \right) \geq \left( \int_{M'} \varrho' dM' \right)^2 = \left( \int_M \varrho dM \right)^2 = \varrho^2 [\text{vol}(M)]^2 = \varrho^2 \text{vol}(M') \text{vol}(M) = \text{vol}(M') \int_M \varrho^2 dM$ . Then by (3.6),  $|B'|^2 = 0$ . Hence, by Corollary 2.3,  $(M', g', J')$  is Kähler and of constant holomorphic sectional curvature  $\kappa' = \kappa$ .

Letting  $p = 1$  in (1.4)–(1.9) gives

$$(3.7) \quad a_{0,1} = 2n \int_M dM = 2n \text{vol}(M),$$

$$(3.8) \quad a_{1,1} = \frac{n-3}{3} \int_M \varrho dM,$$

$$\begin{aligned}
(3.9) \quad a_{2,1} &= \frac{1}{180} \int_M \left[ (5n-30)\varrho^2 + (-2n+90)|Rc|^2 \right. \\
&\quad \left. + (2n-15)|R|^2 \right] dM.
\end{aligned}$$

In dimension  $2n \geq 16$ , substitute (2.5) in (3.9). Then, since  $E = 0$  and  $a_{2,1} = a'_{2,1}$ , we have

$$\begin{aligned}
& \int_M \left[ \frac{(5n^3 - 26n^2 + 12n + 60)\varrho^2 + (6n - 45)\varrho\varrho^*}{n(n+1)} + (2n - 15)|B|^2 \right] dM \\
&= \int_M \left[ \frac{(5n^3 - 26n^2 + 18n + 15)\varrho^2}{n(n+1)} + (2n - 15)|B|^2 \right] dM \\
(3.10) \quad &= \int_{M'} \left[ \frac{(5n^3 - 26n^2 + 12n + 60)\varrho'^2 + (6n - 45)\varrho'\varrho'^*}{n(n+1)} \right. \\
&\quad \left. + (2n - 15)|B'|^2 \right] dM' \\
&\geq \int_{M'} \left[ \frac{(5n^3 - 26n^2 + 18n + 15)\varrho'^2}{n(n+1)} + (2n - 15)|B'|^2 \right] dM'.
\end{aligned}$$

The result then follows in a similar way as in the case  $p = 0$ .

Letting  $p = 2$  in (1.4)–(1.9) gives

$$(3.11) \quad a_{0,2} = (2n^2 - n) \int_M dM = (2n^2 - n) \text{vol}(M),$$

$$(3.12) \quad a_{1,2} = \frac{2n^2 - 13n + 12}{6} \int_M \varrho dM,$$

$$\begin{aligned}
(3.13) \quad a_{2,2} &= \frac{1}{360} \int_M \left[ (10n^2 - 125n + 300)\varrho^2 \right. \\
&\quad \left. + (-4n^2 + 362n - 1080)|Rc|^2 \right. \\
&\quad \left. + (4n^2 - 62n + 240)|R|^2 \right] dM.
\end{aligned}$$

In dimension  $2n = 4, 6, 8, 14$  or  $2n \geq 18$ , substitute (2.5) in (3.13). Then, since  $E = 0$  and  $a_{2,2} = a'_{2,2}$ , we have

$$\begin{aligned}
& \int_M \left[ \frac{(10n^4 - 117n^3 + 350n^2 + 3n - 780)\varrho^2 + (12n^2 - 186n + 720)\varrho\varrho^*}{n(n+1)} \right. \\
&\quad \left. + (4n^2 - 62n + 240)|B|^2 \right] dM
\end{aligned}$$

$$\begin{aligned}
&= \int_M \left[ \frac{(10n^4 - 117n^3 + 362n^2 - 183n - 60)\varrho^2}{n(n+1)} \right. \\
(3.14) \quad &\quad \left. + (4n^2 - 62n + 240)|B|^2 \right] dM \\
&= \int_{M'} \left[ \frac{(10n^4 - 117n^3 + 350n^2 + 3n - 780)\varrho'^2 + (12n^2 - 186n + 720)\varrho'\varrho'^*}{n(n+1)} \right. \\
&\quad \left. + (4n^2 - 62n + 240)|B'|^2 \right] dM' \\
&\geq \int_{M'} \left[ \frac{(10n^4 - 117n^3 + 362n^2 - 183n - 60)\varrho'^2}{n(n+1)} \right. \\
&\quad \left. + (4n^2 - 62n + 240)|B'|^2 \right] dM'.
\end{aligned}$$

The result then follows in a similar way as in the case  $p = 0$ .

If  $(M', g')$  is almost Hermitian Einstein of nonzero scalar curvature  $\varrho'$  with  $\varrho' = \varrho'^*$ , then we have equality in (3.6), (3.10) and (3.14) and the results follow in the same way.

**Corollary 3.3.** *If  $(\mathbb{C}P^n, g_0, J_0)$  is complex projective space with the Fubini–Study metric,  $(M', g', J')$  a compact almost Hermitian Einstein manifold of either positive scalar curvature  $\varrho'$  with  $\varrho' \leq \varrho'^*$  or nonzero scalar curvature  $\varrho'$  with  $\varrho' = \varrho'^*$ , and  $\text{Spec}^p(M', g', J') = \text{Spec}^p(\mathbb{C}P^n, g_0, J_0)$  for  $p = 0, 1$  or  $2$ , then  $(M', g', J')$  is Kähler and holomorphically isometric to  $(\mathbb{C}P^n, g_0, J_0)$  in the following cases:  $p = 0$  and  $2n \geq 2$ , hence  $(\mathbb{C}P^n, g_0, J_0)$  is characterized by the spectrum in every dimension in these classes of manifolds;  $p = 1$  and  $2n = 2$  or  $2n \geq 16$ ;  $p = 2$  and  $2n = 2, 4, 6, 8, 14$  or  $2n \geq 18$ .*

PROOF. Since  $(\mathbb{C}P^n, g_0, J_0)$  is the only Kähler manifold with a metric of positive constant holomorphic sectional curvature  $\kappa$ , then  $(M', g', J')$  is Kähler with constant holomorphic sectional curvature  $\kappa$  and  $(M', g', J')$  and  $(\mathbb{C}P^n, g_0, J_0)$  are holomorphically isometric.

**Corollary 3.4.** *If  $(M, g, J)$  is a compact Kähler manifold of constant holomorphic sectional curvature  $\kappa$  and  $(M', g', J')$  is either compact almost Kähler Einstein with positive scalar curvature or compact nearly Kähler*

*Einstein with negative scalar curvature and  $\text{Spec}^p(M, g) = \text{Spec}^p(M', g')$  for  $p = 0, 1$  or  $2$ , then the conclusion of Theorem 3.2b holds.*

PROOF. The result follows immediately from Lemma 3.1 and Theorem 3.2.

We note here that a compact almost Kähler Einstein manifold of non-negative scalar curvature is Kähler [9].

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