

On weakly Ricci symmetric spaces

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Abstract. First we prove the existence of a weakly Ricci symmetric space by an example. Next we study some properties of a weakly Ricci symmetric space. Among others it is proved that a conformally flat weakly Ricci symmetric space is of almost constant curvature and if the Ricci tensor of the space is cyclic then the sum of the associated 1-forms is zero. Finally it is proved that a conformally flat weakly Ricci symmetric space-time is the Robertson–Walker space-time.

1. Preliminaries

The notions of weakly symmetric and weakly Ricci symmetric spaces were introduced by L. TAMÁSSY and T. Q. BINH in [1], [2]. A non-flat Riemannian space V^n ($n > 2$) is called weakly symmetric if the curvature tensor R_{hijk} satisfies the condition:

$$(1) \quad R_{hijk,\ell} = a_\ell R_{hijk} + b_h R_{lij k} + c_i R_{h\ell j k} + d_j R_{hi\ell k} + e_k R_{hij\ell}$$

where a, b, c, d, e are 1-forms (non-zero simultaneously) “,” in (1) denotes covariant differentiation with respect to the metric tensor. a, b, c, d, e are called the associated 1-forms of the space. This space is denoted by $(WS)_n$. Such a space have been studied by M. PRVANOVIĆ [3], T. Q. BINH [4], U. C. DE and S. BANDYOPADHYAY [5] and others.

A non-flat Riemannian space is called weakly Ricci symmetric and denoted by $(WRS)_n$ if the Ricci tensor is non-zero and satisfies the condition:

$$(2) \quad R_{ij,k} = a_k R_{ij} + b_i R_{kj} + c_j R_{ik}.$$

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where a, b, c , are 1-forms (non-zero simultaneously). If in (1) a_ℓ is replaced by $2a_\ell$ and e_k is replaced by a_k then the space is called a generalized pseudo symmetric space introduced by CHAKI [6], and if in (2) a_k is replaced by $2a_k$ then the space is called a generalized pseudo Ricci symmetric space introduced by CHAKI and KOLEY [7]. So the defining conditions of $(WS)_n$ and $(WRS)_n$ are little weaker than that of generalized pseudo symmetric and generalized pseudo Ricci symmetric spaces respectively. But weakly symmetric and weakly Ricci symmetric spaces are different from that of pseudo symmetric and Ricci pseudo symmetric spaces in the sense of R. DESZCZ [8].

At the study of a $(WRS)_n$ the vector

$$\lambda_i = b_i - c_i$$

plays an important role.

In Section 2 we give a concrete example of a $(WRS)_n$. In Sections 3 and 4 we study $(WRS)_n$ of definite metric and with $\lambda \neq 0$. In Section 3 A) we show that the scalar curvature R does not vanish and $R_{ij} = R\theta_i\theta_j$ where θ is a unit vector, and B) we study the orthogonality of $R_{,h}$ and θ_h . In Section 4 we consider conformally flat $(WRS)_n$ ($n > 3$) and show that the space is of almost constant curvature. In Section 5 we study $(WRS)_n$ of cyclic Ricci tensor. Finally we prove that a conformally flat weakly Ricci symmetric space-time is the Robertson–Walker space-time.

2. Example of $(WRS)_n$

Let each Latin index run over $1, 2, \dots, n$ and each Greek index over $2, 3, \dots, n-1$. We define the metric g in R^n ($n \geq 4$) by the formula (see [9])

$$(3) \quad ds^2 = \varphi(dx^1)^2 + k_{\alpha\beta}dx^\alpha dx^\beta + 2dx^1 dx^n$$

where $[k_{\alpha\beta}]$ is a symmetric and non-singular matrix consisting of constant entries, and φ is a function independent of x^n .

In the metric considered, the only non-vanishing components of Christoffel symbols and Ricci tensor R_{ij} are (see [9])

$$\{1^\beta{}_1\} = (-1/2)k^{\beta\alpha}\varphi_{,\alpha}, \quad \{1^n{}_1\} = (1/2)\varphi_{,1}, \quad \{1^n{}_\alpha\} = (1/2)\varphi_{,\alpha}$$

and

$$(4) \quad R_{11} = (1/2)k^{\alpha\beta}\varphi_{.\alpha\beta}$$

where $(.)$ denotes partial differentiation and $[k^{\alpha\beta}]$ is the inverse matrix of $[k_{\alpha\beta}]$.

Here we assume that $k_{\alpha\beta} = \delta_{\alpha\beta}$ and $\varphi = \delta_{\alpha\beta} x^\alpha x^\beta e^{2x^1}$. In this case φ reduces to

$$\varphi = \sum_{\alpha=2}^{n-1} x^\alpha x^\alpha e^{2x^1}.$$

Hence $\varphi_{.\alpha\beta} = 2\delta_{\alpha\beta}e^{2x^1}$. Then it follows from (4) that the only non-zero components of R_{ij} and $R_{ij,k}$ are R_{11} and $R_{11,1}$ respectively, where

$$(5) \quad R_{11} = (1/2) \sum_{\alpha=1}^{n-1} k^{\alpha\alpha} \varphi_{.\alpha\alpha} = (1/2)(n-2)2 \cdot e^{2x^1} = (n-2)e^{2x^1}$$

$$R_{11,1} = 2(n-2)e^{2x^1}.$$

So neither the Ricci tensor nor its covariant derivative vanish.

We claim that R^n ($n > 3$) with the given metric g is a $(WRS)_n$.

To verify the relation (2) it is sufficient to check the followings:

- (A) $R_{11,1} = a_1 R_{11} + b_1 R_{11} + c_1 R_{11},$
- (B) $R_{11,k} = a_k R_{11} + b_1 R_{k1} + c_1 R_{k1}, \quad k \neq 1$
- (C) $R_{1j,1} = a_1 R_{1j} + b_1 R_{1j} + c_j R_{11}, \quad j \neq 1$
- (D) $R_{i1,1} = a_1 R_{i1} + b_i R_{11} + c_1 R_{i1}, \quad i \neq 1$
- (E) $R_{ij,1} = a_1 R_{ij} + b_i R_{1j} + c_j R_{i1}, \quad i, j \neq 1$
- (F) $R_{ij,k} = a_k R_{ij} + b_i R_{kj} + c_j R_{ik}, \quad i, j \neq 1.$

In virtue of (5), (A) holds iff

$$(6) \quad 2 = a_1(x) + b_1(x) + c_1(x),$$

(B) holds iff

$$(7) \quad 0 = a_k(n-2)e^{2x^1}, \quad k \neq 1 \iff a_2 = \dots = a_n = 0.$$

Similarly (C) and (D) yield

$$(8) \quad \begin{aligned} b_2 = \dots\dots\dots &= b_n = 0, \\ c_2 = \dots\dots\dots &= c_n = 0. \end{aligned}$$

These mean that (2) can be satisfied by a number of $a(x), b(x), c(x)$, namely by those which fulfil (6), (7), (8).

Thus we have the following

Theorem 1. R^n with the metric g , defined by (3) forms a weakly Ricci symmetric space.

Remark. Here the metric defined by (3) is indefinite and $g^{11} = 0$. Therefore $R = g^{ij}R_{ij} = g^{11}R_{11} = 0$.

3. $(WRS)_n$ of definite metric

We assume that $\text{Det } |g_{ij}| \neq 0$ (the space is definite) and $\lambda \neq 0$.

A) $R_{ij,h} - R_{ji,h} = 0$ yields

$$(9) \quad \lambda_j R_{ih} = \lambda_i R_{hj}, \quad \lambda_i = b_i - c_i.$$

Transvecting this by g^{ih} we get

$$(10) \quad \lambda_j R = \lambda^h R_{hj}.$$

Transvecting (9) again by λ^i we obtain

$$(11) \quad \begin{aligned} \lambda^i \lambda_j R_{ih} &= \lambda^i \lambda_i R_{hj} = \|\lambda\|^2 R_{hj}, \quad \text{and hence} \\ R_{hj} &= \frac{1}{\|\lambda\|^2} \lambda_j \lambda^i R_{ih} \stackrel{(10)}{=} \frac{\lambda_j \lambda_h}{\|\lambda\|^2} R = R \theta_j \theta_h, \end{aligned}$$

where $\theta_h = \frac{\lambda_h}{\|\lambda\|^2}$ is a unit vector. From (11) we conclude that $R \neq 0$, for $R = 0$ implies $R_{hj} = 0$ which is inadmissible by the definition of $(WRS)_n$.

B) From (2) we obtain

$$R_{ij,k} - R_{ik,j} = T_k R_{ij} - T_j R_{ik}, \quad \text{where } T_j = a_j - c_j.$$

Then from (11) we have

$$(12) \quad R_{ij,k} - R_{ik,j} = R(T_k\theta_i\theta_j - T_j\theta_i\theta_k).$$

It follows from the 2nd Bianchi identity that $R^j_{k,j} = (1/2)R_{,k}$ ([10, p. 82]). Transvecting the equation (12) by g^{ij} and using the above relation we obtain

$$(13) \quad R_{,k} = 2R(T_k - (\theta^j T_j)\theta_k).$$

Now transvecting again by θ^k , we have

$$(14) \quad \theta^k(R_{,k}) = 2R(\theta^k T_k - \theta^j T_j) = 0.$$

Again transvecting (13) with T^k we get

$$(15) \quad T^k R_{,k} = 2R(T - P^2), \quad \text{where } T = T^k T_k, \quad P = \theta^j T_j.$$

Thus from (11), (14) and (15) we have the following

Theorem 2. *In a $(WRS)_n$ of definite metric, if $b_i - c_i \neq 0$, then*

A) *the scalar curvature R is non-zero and the Ricci tensor R_{ij} is of the form*

$$R_{ij} = R\theta_i\theta_j$$

B) *$R_{,k}$ is orthogonal to θ_k but it is not orthogonal to T_k except $P^2=T$.*

4. Conformally flat $(WRS)_n$ ($n > 3$) of definite metric

For a conformally flat Riemannian space V^n we have ([10, p. 92])

$$R_{hijk} = \frac{1}{n-2}(g_{ij}R_{hk} - g_{ik}R_{hj} + g_{hk}R_{ij} - g_{hj}R_{ik}) - \frac{R}{(n-1)(n-2)}(g_{hk}g_{ij} - g_{hj}g_{ik}).$$

Let us suppose now that this V^n is a $(WRS)_n$, and $\lambda \neq 0$. Then in virtue of (11) we have

$$(16) \quad R_{hijk} = a(g_{hk}g_{ij} - g_{hj}g_{ik}) + b(g_{hj}\theta_i\theta_k - g_{hk}\theta_i\theta_j - g_{ij}\theta_h\theta_k + g_{ik}\theta_h\theta_j),$$

where $a = -\frac{R}{(n-1)(n-2)}$ and $b = -\frac{R}{(n-2)}$.

D. SMARANDA [11] calls a Riemannian space whose curvature tensor satisfies (16), a space of almost constant curvature. Thus we can state the following:

Theorem 3. *In a conformally flat $(WRS)_n$ ($n > 3$) of definite metric, if $\lambda \neq 0$, then the space is of almost constant curvature.*

Remark. The questions similar to that of Theorem 2 and Theorem 3 under different conditions were investigated by CHAKI and KOLEY [7].

5. $(WRS)_n$ with cyclic Ricci tensor

A Riemannian space is said to have cyclic Ricci tensor if the Ricci tensor satisfies

$$(17) \quad R_{ij,k} + R_{jk,i} + R_{ki,j} = 0.$$

Using (2) in (17) we obtain

$$(18) \quad \theta_k R_{ij} + \theta_i R_{kj} + \theta_j R_{ik} = 0,$$

where $\theta_i = a_i + b_i + c_i$.

Now we state

WALKER's Lemma [12]: *If a_{ij} , b_i are numbers satisfying $a_{ij} = a_{ji}$, $a_{ij}b_k + a_{jk}b_i + a_{ki}b_j = 0$ for $i, j, k = 1, 2, \dots, n$, then either all the a_{ij} are zero or all the b_i are zero.*

Hence by the above lemma we get from (18) that either $\theta_i = 0$ or $R_{ij} = 0$. But by definition of $(WRS)_n$, $R_{ij} \neq 0$. Therefore $\theta_i = 0$. Thus we can state the following.

Theorem 4. *If a $(WRS)_n$ satisfies cyclic Ricci tensor, then the sum of the associated 1-forms is zero.*

6. Results concerning the warped product

In an earlier paper [13] the first author and B. K. DE proved that in a conformally flat generalized pseudo Ricci symmetric space the vector $\lambda_i = b_i - c_i$ is a proper concircular vector field. A weakly Ricci symmetric space is almost the same as a generalized pseudo Ricci symmetric space. Thus it is easy to see that a conformally flat $(WRS)_n$ admits a concircular vector field. K. YANO [14] proved that in order that a Riemannian space admits a concircular vector field, it is necessary and sufficient that there exists a coordinate system with respect to which the fundamental quadratic

differential form may be written in the form

$$ds^2 = (dx^1)^2 + f(x^1) g_{\alpha\beta}^* dx^\alpha dx^\beta,$$

where $g_{\alpha\beta}^* = g_{\alpha\beta}^*(x^\gamma)$ are the functions of x^γ only ($\alpha, \beta, \gamma = 2, 3, \dots, n$) and f is a function of x^1 only.

Similarly we can prove that a Lorentzian space with the metric of signature $(-+++)$ admits a concircular vector field if and only if there exists a coordinate system with respect to which the fundamental quadratic differential form may be written in the form

$$ds^2 = -(dx^1)^2 + f(x^1) g_{jk}^* dx^j dx^k.$$

In this section we consider a $(WRS)_n$ space-time. By a space-time, we will mean a 4-dimensional space endowed with Lorentz metric of signature $(-+++)$. Since a conformally flat $(WRS)_n$ admits a concircular vector field, therefore the conformally flat $(WRS)_n$ can be expressed as the warped product $I \times_f M^*$ where I is an open interval of R and (M, g^*) is a 3-dimensional Riemannian space.

In a conformally flat space we have [10]

$$(19) \quad R_{ij,k} - R_{ik,j} = \frac{1}{2(n-1)} (R_{,k} g_{ij} - R_{,j} g_{ik}).$$

A. GEBAROWSKI [15] proved that a warped product $I \times_f M^*$ satisfies (19) if and only if M^* is an Einstein space. Thus if a $(WRS)_n$ space-time is conformally flat, it must be a warped product $I \times_f M^*$, where M^* is a 3-dimensional Einstein space. It is known ([10]) that a 3-dimensional Einstein space is a space of constant curvature. Hence a conformally flat $(WRS)_n$ space-time is the warped product $I \times_f M^*$, where M^* is a space of constant curvature. But such a warped product is Robertson–Walker space-time [16].

Thus we have the following

Theorem 5. *A conformally flat weakly Ricci symmetric space-time is the Robertson–Walker space-time.*

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